#### QUANTUM STRUCTURE OF SPACETIME AND GRAVITY

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# MAXWELL-WEYL GAUGE THEORY OF GRAVITY

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#### **MOTIVATION**

- 1. Maxwell group
- 2. Summary of Literature
- 3. Influence area of Maxwell group

# What is the Maxwell Algebra?

- The Maxwell algebra is a non-central extension of the Poincare algebra, in which the momentum generators no longer commute, but satisfy  $[P_a, P_b] = Z_{ab}$ . The charges  $Z_{ab}$  commute with the momenta, and transform tensorially under the action of the angular momentum generators.
- If one constructs an action for a massive particle, invariant under these symmetries, one finds that it satisfies the equations of motion of a charged particle interacting with a constant electromagnetic field via the Lorentz force.

G.W. Gibbons, J. Gomis and C. N. Pope, PHYSICAL REVIEW D 82, 065002 (2010)

$$[P_a, P_b] = 0 \rightarrow [\tilde{P}_a, \tilde{P}_b] = iZ_{ab}$$

### Maxwell Algebra

 If we enlarge the Poincare algebra by six additional tensorial abelian generators as showed below then we get Maxwell algebra.

#### Maxwell algebra

$$\begin{split} \left[ M_{ab}, M_{cd} \right] &= i \left( \eta_{ad} M_{bc} + \eta_{bc} M_{ad} - \eta_{ac} M_{bd} - \eta_{bd} M_{ac} \right) \\ \left[ M_{ab}, \Pi_c \right] &= i \left( \eta_{bc} \Pi_a - \eta_{ac} \Pi_b \right) \\ \left[ \Pi_a, \Pi_b \right] &= i F_{ab} \\ \left[ M_{ab}, F_{cd} \right] &= i \left( \eta_{ad} F_{bc} + \eta_{bc} F_{ad} - \eta_{ac} F_{bd} - \eta_{bd} F_{ac} \right) \\ \left[ F_{ab}, F_{cd} \right] &= 0 \\ \left[ F_{ab}, \Pi_c \right] &= 0 \end{split}$$

New momentum op.

$$\begin{split} & \Pi_{a} = i \left( \partial_{a} + e A_{a} \left( x \right) \right) & \longrightarrow \left( i \gamma^{a} \Pi_{a} - m \right) \psi \left( x, f \right) = 0 \\ & A_{a} \left( x \right) = \frac{1}{2} f_{a}^{b} x_{b} & \longrightarrow \left[ f_{ab} = \partial_{a} A_{b} \left( x \right) - \partial_{b} A_{a} \left( x \right) \right] \end{split}$$

Where,

 $A_a$ : Electromagnetic potential, e: Electric charge  $f_{ab}$ : Electromagnetic field tensor

This antisymmetric generator  $F_{[ab]}$  can be used to describe the motion of a relativistic particle in a constant electromagnetic field.

If we select  $F_{[ab]} = 0$  then we get well-known Poincare algebra which describes flat Minkowski spacetime.

H. Bacry, P. Combe, and J.L. Richard, Nuovo Cim. 67, (1970), 267 R. Schrader, Fortsch.Phys. 20, (1972), 701-734

# Maxwell Algebra

In 2005 D. V. Soroka, V. A. Soroka shows us the Maxwell algebra can be found with new additional tensorial coordinate  $\theta^{ab}$  and corresponding tensorial derivative  $\partial_{ab} = \frac{\partial}{\partial ab}$ .

#### Maxwell group

$$\begin{bmatrix} M_{ab}, M_{cd} \end{bmatrix} = i \left( \eta_{ad} M_{bc} + \eta_{bc} M_{ad} - \eta_{ac} M_{bd} - \eta_{bd} M_{ac} \right)$$
$$\begin{bmatrix} M_{ab}, P_{c} \end{bmatrix} = i \left( \eta_{bc} P_{a} - \eta_{ac} P_{b} \right)$$

$$[P_a, P_b] = iZ_{ab}$$

$$\left[M_{ab},Z_{cd}\right]\!=\!i\!\left(\eta_{ad}Z_{bc}+\eta_{bc}Z_{ad}-\eta_{ac}Z_{bd}-\eta_{bd}Z_{ac}\right)$$

$$[Z_{ab}, Z_{cd}] = 0$$

$$[Z_{ab}, P_{c}] = 0$$

Differential realisation of generators:

$$P_{a} = i(\partial_{a} - \frac{1}{2} X^{b} \partial_{ab})$$

$$Z_{ab} = i\partial_{ab}$$

$$|\mathsf{M}_{\mathsf{a}\mathsf{b}} = \mathsf{i}(\mathsf{x}_{\mathsf{a}}\partial_{\mathsf{b}} - \mathsf{x}_{\mathsf{b}}\partial_{\mathsf{a}}) + 2\mathsf{i}(\theta_{\mathsf{a}}{}^{\mathsf{c}}\partial_{\mathsf{b}\mathsf{c}} - \theta_{\mathsf{b}}{}^{\mathsf{c}}\partial_{\mathsf{a}\mathsf{c}})$$

D. V. Soroka, V. A. Soroka, Physics Letters B 607 (2005) 302-305.

#### Some relations of tensorial coordinates and its derivatives:

$$\left| \partial_{ab} \theta^{cd} = \frac{1}{2} \left( \delta_a^c \delta_b^d - \delta_b^c \delta_a^d \right) \right| \left[ \partial_{ab} x^c = 0 \right] \left[ \left[ \partial_{ab}, \partial_c \right] = 0 \right] \left[ \left[ \partial_{ab}, \partial_{cd} \right] = 0$$

$$\partial_{ab} x^c = 0$$

$$\left[\partial_{ab},\partial_{c}\right]=0$$

$$\left[\left(\partial_{ab},\partial_{cd}\right)\right]=0$$

# A Brief Summary of Litarature

#### Gauge Theories of Some Important Groups

Lorentz group

Poincare group

Weyl group

De Sitter group

Affine group

J.M. Charap and W.

#### Gauge Theories of Maxwell Groups

Simple Maxwell Group

J.A. Azcarraga, K. D.V. Soroka, V.A. Kamimura, J. Lukierski

Semisimple Maxwell Group

Soroka - 2011

Maxwell-Weyl Group

O. Cebecioğlu, S. Kibaroğlu -2014

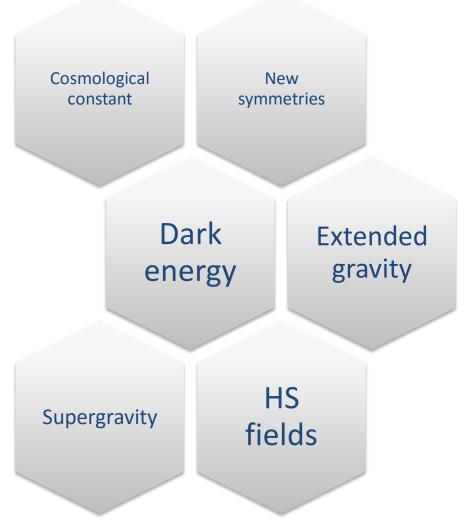
AdS-Maxwell Group

R. Durka, J. Kowalski-Glikman, M. Szczachor – 2011

Maxwell-Affine Group

O. Cebecioğlu, S. Kibaroğlu -2015

# Influence Area of Maxwell Group



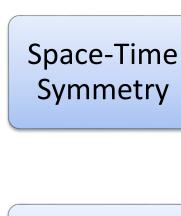
### **GENERAL OVERVIEW**

1. Aim and method

### Aim

Our fundamental aim is to going beyond the basic gravity.

### Method





Extension with antisymmetric tensor generator



Extended
Space-Time
(Maxwell)
Symmetry



Supersymmetric extension of the extended symmetry



Extended Gravity (Maxwell Gravity)



#### Gauge Theory

• Differential Geometry



Super Maxwell Gravity

### MAXWELL-WEYL GROUP (MW(1,3))

- 1. Tensor extension of Weyl algebra
- 2. Gauge theory of Maxwell-Weyl group

#### Weyl algebra:

$$\begin{split} & \left[ \left[ M_{ab}, M_{cd} \right] = i \left( \eta_{ad} M_{bc} + \eta_{bc} M_{ad} - \eta_{ac} M_{bd} - \eta_{bd} M_{ac} \right) \\ & \left[ M_{ab}, P_c \right] = i \left( \eta_{bc} P_a - \eta_{ac} P_b \right) \end{split}$$

$$[P_a, P_b] = 0$$

$$[D,D]=0$$

$$[P_a,D]=iP_a$$

$$[M_{ab},D]=0$$

Weyl algebra contains three generators  $(M_{ab}, P_a, D)$ . These genrators correspond to Lorentz, momentum and scale symmetry respectively.

Differential realisation of the generators:

$$\begin{aligned} & P_{a} = i\partial_{a} \\ & D = i(x \cdot \partial) \\ & M_{ab} = i(x_{a}\partial_{b} - x_{b}\partial_{a}) \end{aligned}$$

#### Maxwell-Weyl algebra:

$$\begin{split} & \left[ M_{ab}, M_{cd} \right] = i \left( \eta_{ad} M_{bc} + \eta_{bc} M_{ad} - \eta_{ac} M_{bd} - \eta_{bd} M_{ac} \right) \\ & \left[ M_{ab}, P_c \right] = i \left( \eta_{bc} P_a - \eta_{ac} P_b \right) \\ & \left[ P_a, P_b \right] = i Z_{ab} \\ & \left[ D, D \right] = 0 \\ & \left[ P_a, D \right] = i P_a \\ & \left[ M_{ab}, D \right] = 0 \\ & \left[ M_{ab}, Z_{cd} \right] = i \left( \eta_{ad} Z_{bc} + \eta_{bc} Z_{ad} - \eta_{ac} Z_{bd} - \eta_{bd} Z_{ac} \right) \\ & \left[ Z_{ab}, Z_{cd} \right] = 0 \\ & \left[ Z_{ab}, P_c \right] = 0 \\ & \left[ Z_{ab}, D \right] = 2i Z_{ab} \end{split}$$

Weyl algebra can be extended as in the box.

This algebra satisfy all Jacobi Identities.

Where  $Z_{ab}$  are antisymmetric generators.

One can find this algebra with different notation in following paper:

S. Bonanos, J. Gomis, K. Kamimura, and J. Lukierski, Journal of Math. Phys. 51, 102301 (2010)

In order to find differential realisation of generators, we use following coset transformation;

$$K(x', \theta', \sigma')h(\omega) = g'(a, \varepsilon, \lambda, u)K(x, \theta, \sigma)$$

#### Where,

$$g\!\left(x, \! \! \boldsymbol{\theta}, \boldsymbol{\sigma}, \boldsymbol{\omega}\right) = e^{ix(x) \cdot P} e^{i\theta\!\left(x\right) \cdot Z} e^{i\sigma\!\left(x\right) \cdot D} e^{-\frac{i}{2}\boldsymbol{\omega}\!\left(x\right) \cdot M}$$

$$g\big(a, \underline{\epsilon}, \lambda, u\big) = e^{ia \cdot P} e^{i\underline{\epsilon} \cdot Z} e^{i\lambda \cdot D} e^{-\frac{i}{2}u \cdot M}$$

$$K(x, \theta, \sigma) = e^{ix(x) \cdot P} e^{i\theta(x) \cdot Z} e^{i\sigma(x) \cdot D}$$



$$\begin{split} \delta x^{a} &= a^{a} + u^{a}_{b} x^{b} + \lambda x^{a} \\ \delta \theta^{ab} &= \epsilon^{ab} + u^{[a}_{c} \theta^{cb]} + 2\lambda \theta^{ab} - \frac{1}{4} a^{[a} x^{b]} \\ \delta \sigma &= \lambda \\ \delta \omega^{ab} &= u^{ab} \end{split}$$

Trasformation of scalar field is defined by;

$$\Phi'\left(\mathbf{x}^{\mathsf{a}}, \boldsymbol{\theta}^{\mathsf{ab}}\right) = \Phi\left(\mathbf{x}^{\mathsf{a}} - \delta\mathbf{x}^{\mathsf{a}}, \boldsymbol{\theta}^{\mathsf{ab}} - \delta\boldsymbol{\theta}^{\mathsf{ab}}\right)$$

If we use  $\delta x$  and  $\delta \theta$  on the trasformation law of scalar field and compare following equation;

$$\delta\Phi\left(x, \theta\right) = \left(ia \cdot P + \frac{i}{i\epsilon} \cdot Z + i\lambda D - \frac{i}{2}u \cdot M\right)\Phi\left(x, \theta\right)$$

#### Then we get;

$$\begin{aligned} & P_{a} = i(\partial_{a} - \frac{1}{2} x^{b} \partial_{ab}) \\ & Z_{ab} = i \partial_{ab} \\ & D = i \left( x^{a} \partial_{a} + 2 \theta^{ab} \partial_{ab} \right) \\ & M_{ab} = i (x_{a} \partial_{b} - x_{b} \partial_{a}) + 2 i (\theta_{a}^{\ c} \partial_{bc} - \theta_{b}^{\ c} \partial_{ac}) \end{aligned}$$

#### Defining gauge field ( A ) by;

$$\begin{split} \boldsymbol{\mathcal{A}} &= \boldsymbol{\mathcal{A}}^{\mathsf{A}} \boldsymbol{X}_{\mathsf{A}} \\ &= e^{\mathsf{a}} \boldsymbol{P}_{\mathsf{a}} + \boldsymbol{B}^{\mathsf{ab}} \boldsymbol{Z}_{\mathsf{ab}} + \chi \boldsymbol{D} - \frac{1}{2} \, \boldsymbol{\omega}^{\mathsf{ab}} \boldsymbol{M}_{\mathsf{ab}} \end{split}$$

Where  $X^A$  represents group generators and  $e^a$ ,  $B^{ab}$ ,  $\chi$ ,  $\omega^{ab}$  are associated gauge fields.

If we going over the space-time indices  $e^a$ ,  $B^{ab}$ ,  $\chi$  and  $\omega^{ab}$  take the following form;

$$e^{a}=e_{\mu}^{a}dx^{\mu},\ \omega^{ab}=\omega_{\mu}^{ab}dx^{\mu},\ \chi=\chi_{\mu}dx^{\mu},\ B^{ab}=B_{\mu}^{ab}dx^{\mu}$$

Also gauge field and its variation can be written by;

$$\begin{split} & \mathcal{A}_{\!\mu} = e^a_{\mu} P_a + B^{ab}_{\mu} Z_{ab} + \chi_{\mu} D - \tfrac{1}{2} \, \omega^{ab}_{\mu} M_{ab} \\ & \delta \mathcal{A}_{\!\mu} = \delta e^a_{\mu} P_a + \delta B^{ab}_{\mu} Z_{ab} + \delta \chi_{\mu} D - \tfrac{1}{2} \, \delta \omega^{ab}_{\mu} M_{ab} \end{split}$$

To find variation of gauge field, we can use the following formula;

$$\delta A_{\mu} = -\partial_{\mu} \zeta - i \left[ A_{\mu}, \zeta \right]$$

Under a local gauge transformation with the values in Maxwell algebra  $\zeta(x)$ ;

$$\begin{split} \zeta\left(x\right) &= \zeta^{\mathsf{A}}\left(x\right)X_{\mathsf{A}} \\ &= y^{\mathsf{a}}\left(x\right)P_{\mathsf{a}} + \phi^{\mathsf{ab}}\left(x\right)Z_{\mathsf{ab}} + \rho(x)D - \frac{1}{2}\tau^{\mathsf{ab}}\left(x\right)M_{\mathsf{ab}} \end{split}$$

#### We get following transformations;

$$\begin{split} \delta e_{\mu}^{a} &= -\partial_{\mu} y^{a} - \omega_{\mu \, c}^{a} y^{c} - \chi_{\mu} y^{a} + e_{\mu}^{c} \tau_{\ c}^{a} + \rho e_{\mu}^{a} \\ \delta B_{\mu}^{ab} &= -\partial_{\mu} \phi^{ab} - \omega_{\mu \, c}^{[a} \phi^{cb]} - 2\chi_{\mu} \phi^{ab} + \tau_{\ c}^{[a} B_{\mu}^{cb]} + 2\rho B_{\mu}^{ab} + \frac{1}{2} e_{\mu}^{[a} y^{b]} \\ \delta \chi_{\mu} &= -\partial_{\mu} \rho \\ \delta \omega_{\mu}^{ab} &= -\partial_{\mu} \tau^{ab} - \omega_{\mu \, c}^{[a} \tau^{cb]} \end{split}$$

The curvature forms can be found with the use of following equations;

$$\mathbb{F} = dA + \frac{i}{2} [A, A]$$

Where:

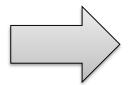
$$\mathcal{F} = \mathcal{F}^{\mathsf{A}} \mathsf{X}_{\mathsf{A}} = \mathsf{F}^{\mathsf{a}} \mathsf{P}_{\mathsf{a}} + \mathsf{F}^{\mathsf{a}\mathsf{b}} \mathsf{Z}_{\mathsf{a}\mathsf{b}} + \mathsf{f} \mathsf{D} - \frac{1}{2} \mathsf{R}^{\mathsf{a}\mathsf{b}} \mathsf{M}_{\mathsf{a}\mathsf{b}}$$

#### One gets;

$$\begin{split} F^{a} &= de^{a} + \omega^{a}_{\ b} \wedge e^{b} + \chi \wedge e^{a} \\ F^{ab} &= dB^{ab} + \omega^{[a}_{\ c} \wedge B^{cb]} + 2\chi \wedge B^{ab} - \frac{1}{2}e^{a} \wedge e^{b} \\ f &= d\chi \\ R^{ab} &= d\omega^{ab} + \omega^{a}_{\ c} \wedge \omega^{cb} \end{split}$$

#### If we going over space-time indices;

$$\begin{split} F^{a} &= \tfrac{1}{2} F^{a}_{\mu\nu} dx^{\mu} \wedge dx^{\nu} \\ F^{ab} &= \tfrac{1}{2} F^{ab}_{\mu\nu} dx^{\mu} \wedge dx^{\nu} \\ f &= \tfrac{1}{2} f_{\mu\nu} dx^{\mu} \wedge dx^{\nu} \\ R^{ab} &= \tfrac{1}{2} R^{ab}_{\mu\nu} dx^{\mu} \wedge dx^{\nu} \end{split}$$



$$\begin{split} & F^a_{\mu\nu} = \partial_{[\mu} e^a_{\nu]} + \omega^a_{[\mu b} e^b_{\nu]} + \chi_{[\mu} e^a_{\nu]} \\ & F^{ab}_{\mu\nu} = \partial_{[\mu} B^{ab}_{\nu]} + \omega^{[a}_{[\mu|c} B^{cb]}_{\nu]} + 2\chi_{[\mu} B^{ab}_{\nu]} - \frac{1}{2} e^a_{[\mu} e^b_{\nu]} \\ & f_{\mu\nu} = \partial_{[\mu} \chi_{\nu]} \\ & R^{ab}_{\mu\nu} = \partial_{[\mu} \omega^{ab}_{\nu]} + \omega^a_{[\mu|c} \omega^{cb}_{\nu]} \end{split}$$

To find the variation of curvatures under local gauge transformation one uses;

$$\delta \mathscr{F}_{\mu\nu} = \mathbf{i} \Big[ \zeta, \mathscr{F}_{\mu\nu} \Big] \Big|$$

#### We get;

$$\begin{split} \delta F^{a}_{\mu\nu} &= -y^{a} f_{\mu\nu} - R^{a}_{\mu\nu} \,_{b} y^{b} + \rho F^{a}_{\mu\nu} + \tau^{a}_{\ b} F^{b}_{\mu\nu} \\ \delta F^{ab}_{\mu\nu} &= -2 \phi^{ab} F_{\mu\nu} + \phi^{[a}_{\ c} R^{cb]}_{\mu\nu} - \frac{1}{2} \, y^{[a} F^{b]}_{\mu\nu} + 2 \rho F^{ab}_{\mu\nu} + \tau^{[a}_{\ c} F^{cb]}_{\mu\nu} \\ \delta f_{\mu\nu} &= 0 \\ \delta R^{ab}_{\mu\nu} &= \tau^{[a}_{\ c} R^{cb]}_{\mu\nu} \end{split}$$

To find an invariant Lagrangian under gauge transformation, we have to take into consideration scale(dilatation) symmetry.

In our case, variation of metric tensor does not vanish. This issue relates to Weyl gauge theory.

$$\delta g_{\mu\nu}\left(x\right) = 2\rho\left(x\right)g_{\mu\nu}\left(x\right)$$

This is major difficulty of constructing of an invariant Lagrangian.

Describing a weyl weight by 'w(f)' then one gets expressions about metric.

$$\begin{aligned} & \left| w \left( g_{\mu\nu} \right) = +2 \\ & w \left( g^{\mu\nu} \right) = -2 \\ & w \left( \sqrt{g_{\mu\nu}} \right) = w \left( e \right) = +4 \end{aligned}$$

$$S_f = \int d^4xe\mathcal{L}_f$$

The free gravitational action requires Weyl weight zero. This condition implies that the Lagrangian density must have  $w(L_f) = -4$ .

Weyl weights of curvatures and gauge fields written as;

$$\begin{aligned} w\left(F_{\mu\nu}^{a}\right) &= 1 \\ w\left(F_{\mu\nu}^{ab}\right) &= 2 \\ w\left(F_{\mu\nu}^{ab}\right) &= 2 \\ w\left(F_{\mu\nu}^{ab}\right) &= 0 \\ w\left(\chi_{\mu\nu}\right) &= 0 \\ w\left(\chi_{\mu\nu$$

One can easily see that the Einstein-Hilbert action does not invariant under the scale transformation.

$$S_{E-H} = \frac{1}{2\kappa} \int d^4x eR$$

To overcome this difficulty, if we multiply it by a compensating scalar field  $\phi$  introduced by Brans-Dicke (1961) and elaborated by Dirac (1973), we can form a Weyl invariant action linear in R.

In our approach we will follow Dirac's idea. The scalar field  $\phi$  with  $w(\phi) = -1$  lets  $\phi^2 R$  be a regular part of  $L_f$ , and hence the combination is invariant under scale transformation.

$$S = \frac{1}{2\kappa} \int d^4x e \phi^2 R$$

C. Brans and R. H. Dicke, Phys. Rev. **124**, 925 (1961) P. A. M. Dirac, Proc. R. Soc. Lon. A. 333, 403-418 (1973)

#### Covariant derivative can be written following form;

$$\mathcal{D}\Phi = \left[ \text{d} + \omega + \text{w} \left( \Phi \right) \chi \right] \Phi$$

#### The curvatures take form;

$$\begin{aligned} & F^{a} = \mathcal{D}e^{a} \\ & F^{ab} = \mathcal{D}B^{ab} - \frac{1}{2}e^{a} \wedge e^{b} \\ & f = d\chi \\ & R^{ab} = \mathcal{D}\omega^{ab} \end{aligned}$$

#### From here one can write Bianchi identities;

$$\begin{split} \mathcal{D}F^{a} &= R^{a}_{\ b} \wedge e^{b} + f \wedge e^{a}, \\ \mathcal{D}F^{ab} &= R^{[a}_{\ c} \wedge B^{cb]} + 2f \wedge B^{ab} - \frac{1}{2}F^{[a} \wedge e^{b]}, \\ \mathcal{D}R^{ab} &= 0, \quad \mathcal{D}f = 0. \end{split}$$

Now we can start to construct Lagrangian.

It is easy to see that following equation has zero weyl weight. We can start to construct Lagrangian with this shifted curvature;

$$\mathcal{J}^{\text{ab}} = R^{\text{ab}} + 2\gamma \phi^2 F^{\text{ab}}$$

With this combination we have an Einstein Lagrangian that involves the curvature scalar linearly. Therefore we consider the following Lagrangian density 4-form as our starting point for the free gravitational part:

$$\begin{split} \mathcal{L}_{\text{f}} &= \frac{1}{2\kappa\gamma} \, J \wedge \,^* J = \frac{1}{4\kappa\gamma} \, \epsilon_{\text{abcd}} J^{\text{ab}} \wedge J^{\text{cd}} \\ &= \frac{1}{4\kappa\gamma} \, \epsilon_{\text{abcd}} R^{\text{ab}} \wedge R^{\text{cd}} + \phi^2 \, \frac{1}{\kappa} \, \epsilon_{\text{abcd}} R^{\text{ab}} \wedge F^{\text{cd}} + \phi^4 \, \frac{\gamma}{\kappa} \, \epsilon_{\text{abcd}} F^{\text{ab}} \wedge F^{\text{cd}} \end{split}$$

Where  $\gamma$  and  $\kappa$  are constants and the first term can be ignored because it is a closed form.

The introduction of a compensating field forces us to add its kinetic term to the Lagrangian. We then get the total action for vacuum as follows:

$$\mathcal{L}_{0} = f \wedge f - \frac{1}{2} \mathcal{D} \phi \wedge {}^{*} \mathcal{D} \phi + \frac{\lambda}{4} \phi^{4} {}^{*} 1$$

Where  $\lambda$  is another constant. Our complete action is the sum of the free gravity action and the vacuum action,

$$S = \int \left( \frac{1}{4\kappa\gamma} \epsilon_{abcd} R^{ab} \wedge R^{cd} + \phi^2 \frac{1}{\kappa} \epsilon_{abcd} R^{ab} \wedge F^{cd} + \phi^4 \frac{\gamma}{\kappa} \epsilon_{abcd} F^{ab} +$$

#### Then we get following equations of motion;

$$\begin{split} \delta_{\omega} S &= 0 \ \, \rightarrow \ \, \mathcal{D} \left( \varphi^2 F^{ab} \right) - \mathcal{J}^{[a}_{\ c} \wedge B^{cb]} = 0 \\ \delta_{e} S &= 0 \ \, \rightarrow \ \, \left[ \begin{matrix} -\varphi^2 \, \frac{1}{\kappa} \, \epsilon_{abcd} \mathcal{J}^{ab} \wedge e^d + \frac{1}{2} \Big[ \, \mathcal{D}_c \varphi^* \mathcal{D} \varphi + \mathcal{D} \varphi \wedge^* \left( e_a \wedge e_c \right) \mathcal{D}^a \varphi \Big] \right] \\ -\frac{1}{4} \Big( f_{cb} e^b \wedge^* f - \frac{1}{2} \, \epsilon_{abcd} f^{ab} e^d \wedge f \Big) + \frac{\lambda}{4} \varphi^4 \,^* e_c \\ \delta_{B} S &= 0 \ \, \rightarrow \ \, \mathcal{D} \left( \varphi^2 \mathcal{J}^{ab} \right) = 0 \\ \delta_{\chi} S &= 0 \ \, \rightarrow \ \, \frac{2\varphi^2}{\kappa} \, \epsilon_{abcd} \mathcal{J}^{ab} \wedge B^{cd} + \frac{1}{2} \, \mathcal{D}^* f + \varphi^* \mathcal{D} \varphi = 0 \\ \delta_{\varphi} S &= 0 \ \, \rightarrow \ \, \frac{2\varphi}{\kappa} \, \epsilon_{abcd} \mathcal{J}^{ab} \wedge F^{cd} + \mathcal{D}^* \mathcal{D} \varphi + \lambda \varphi^3 \,^* 1 = 0 \end{split}$$

One can find following expression thanks to using second equation of previous page;

$$\left| \phi^2 \bigg( \mathcal{J}^{a}_{\phantom{a}b} - \frac{1}{2} \delta^{a}_{\phantom{a}b} \mathcal{J} \bigg) = -\frac{\kappa}{2} \Bigg[ \mathcal{D}^{a} \phi \mathcal{D}_{b} \phi - \frac{1}{2} \delta^{a}_{\phantom{a}b} \bigg( \mathcal{D}^{c} \phi \mathcal{D}_{c} \phi - \frac{\lambda}{2} \phi^4 \bigg) + \bigg( f^{ac} f_{cb} - \frac{1}{4} \delta^{a}_{\phantom{a}b} f^{cd} f_{cd} \bigg) \Bigg] \right|$$

If we swich from tangent space indices to space-time indices then one gets the field equation with a cosmological term depending on the dilaton field;

$$\left|R^{\mu}_{\ \alpha} - \frac{1}{2}\delta^{\mu}_{\ \alpha}R - \frac{3\gamma\phi^2\delta^{\mu}_{\ \alpha}}{} = \frac{2\gamma T\left(B\right)^{\mu}_{\ \alpha} - \frac{\kappa}{2}\phi^{-2}\left[T\left(\phi\right)^{\mu}_{\ \alpha} + T\left(f\right)^{\mu}_{\ \alpha}\right]\right|$$

The energy-momentum tensors:

$$\left|T\left(B\right)^{\mu}_{\phantom{\mu}\alpha}=e^{\mu}_{a}e^{\beta}_{b}\mathcal{D}_{\left[\alpha\right.}B^{ab}_{\beta]}-\frac{1}{2}\delta^{\mu}_{\phantom{\mu}\alpha}\left(e^{\rho}_{a}e^{\sigma}_{b}\mathcal{D}_{\left[\rho\right.}B^{ab}_{\sigma]}\right)\right|$$

$$\left| T \left( \phi \right)^{\mu}_{\ \alpha} = \mathcal{D}^{\mu} \phi \mathcal{D}_{\alpha} \phi - \frac{1}{2} \delta^{\mu}_{\ \alpha} \left( \mathcal{D}^{\gamma} \phi \mathcal{D}_{\gamma} \phi - \frac{\lambda}{2} \phi^{4} \right) \right|$$

O. Cebecioğlu and S. Kibaroğlu PHYSICAL REVIEW D 90, 084053 (2014)

$$\left| \mathsf{T} \left( \mathsf{f} \right)^{\mu}_{\ \alpha} = \mathsf{f}^{\mu\beta} \mathsf{f}_{\beta\alpha} - \frac{1}{4} \delta^{\mu}_{\ \alpha} \mathsf{f}^{\gamma\delta} \mathsf{f}_{\gamma\delta} \right|$$

# Thank you